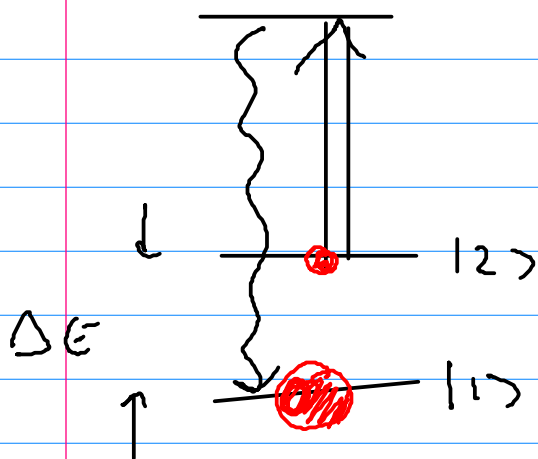
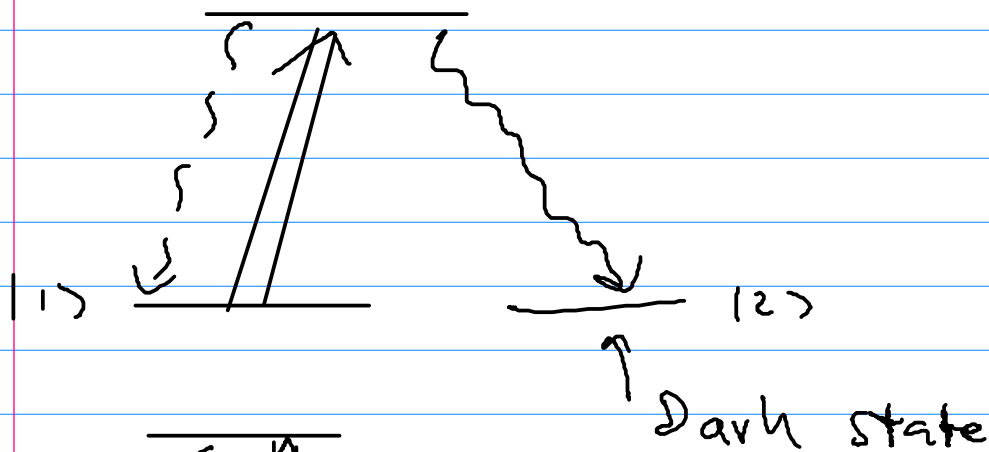


Optical cooling      Sub-Doppler } Temp.  
    Sub-Recoil }

Cooling by optical pumping



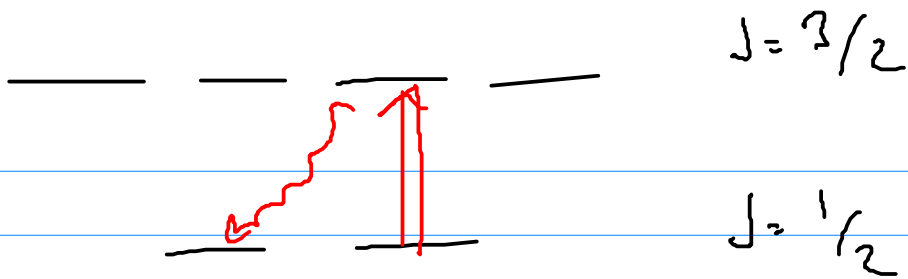
$$\frac{n_2}{n_1} = e^{-\beta \Delta E}$$

$$n_2 \rightarrow 0 \quad T \rightarrow 0$$

Sub-Doppler cooling

New effects in multi-level atoms

"Polarization gradient cooling" is the most famous example



resonances between ground states

width  $\Gamma' = \frac{1}{\tau_p} \sim \frac{\Omega^2}{\delta^2} \Gamma \ll \Gamma$

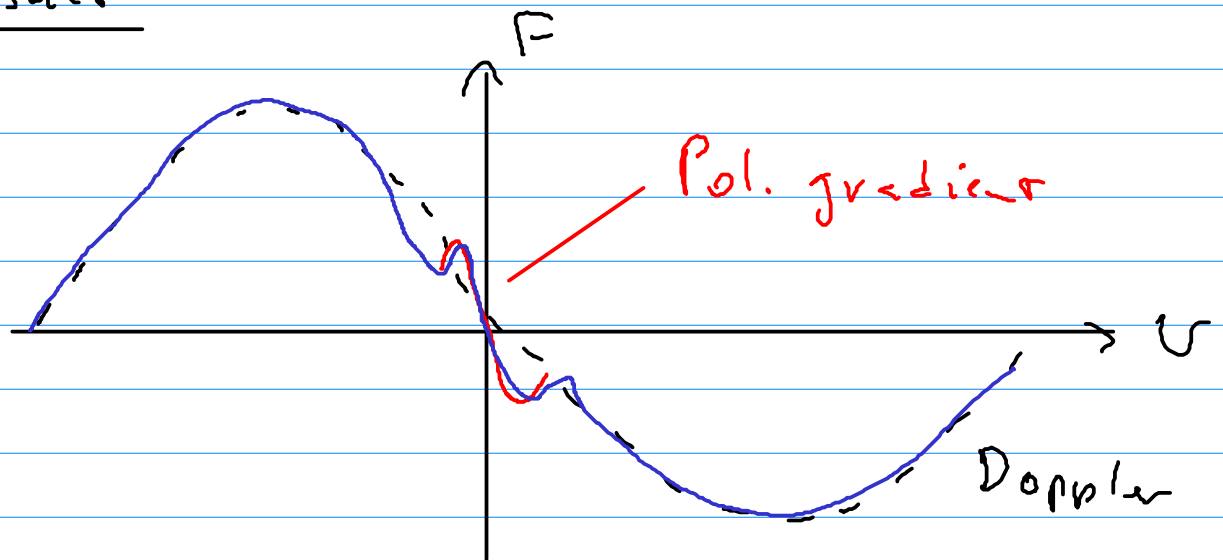
$\Rightarrow$  long relaxation time  $\xrightarrow{\text{optical pumping rate}}$

$\Rightarrow$  possibility of large time lags  
 $\Rightarrow$  cooling

often  $k_B T_{\text{final}} \propto \hbar \Gamma'$  eg. Doppler cooling

(but here  $\Gamma' \frac{\hbar}{\delta}$ )

Result



# Laser cooling below the Doppler limit by polarization gradients: simple theoretical models

J. Dalibard and C. Cohen-Tannoudji

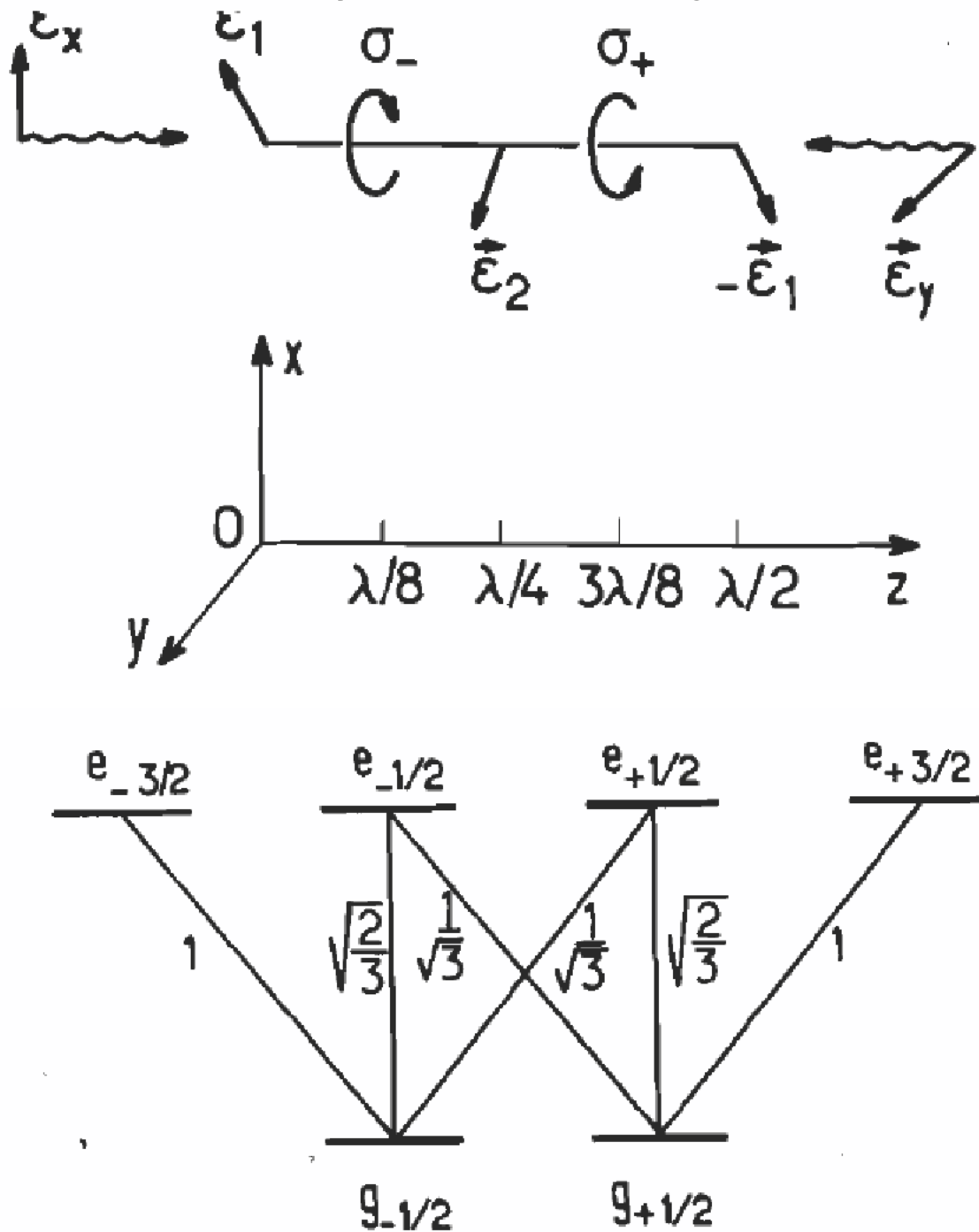


Fig. 2. Atomic level scheme and Clebsch-Gordan coefficients for a  $J_g = 1/2 \leftrightarrow J_e = 3/2$  transition.

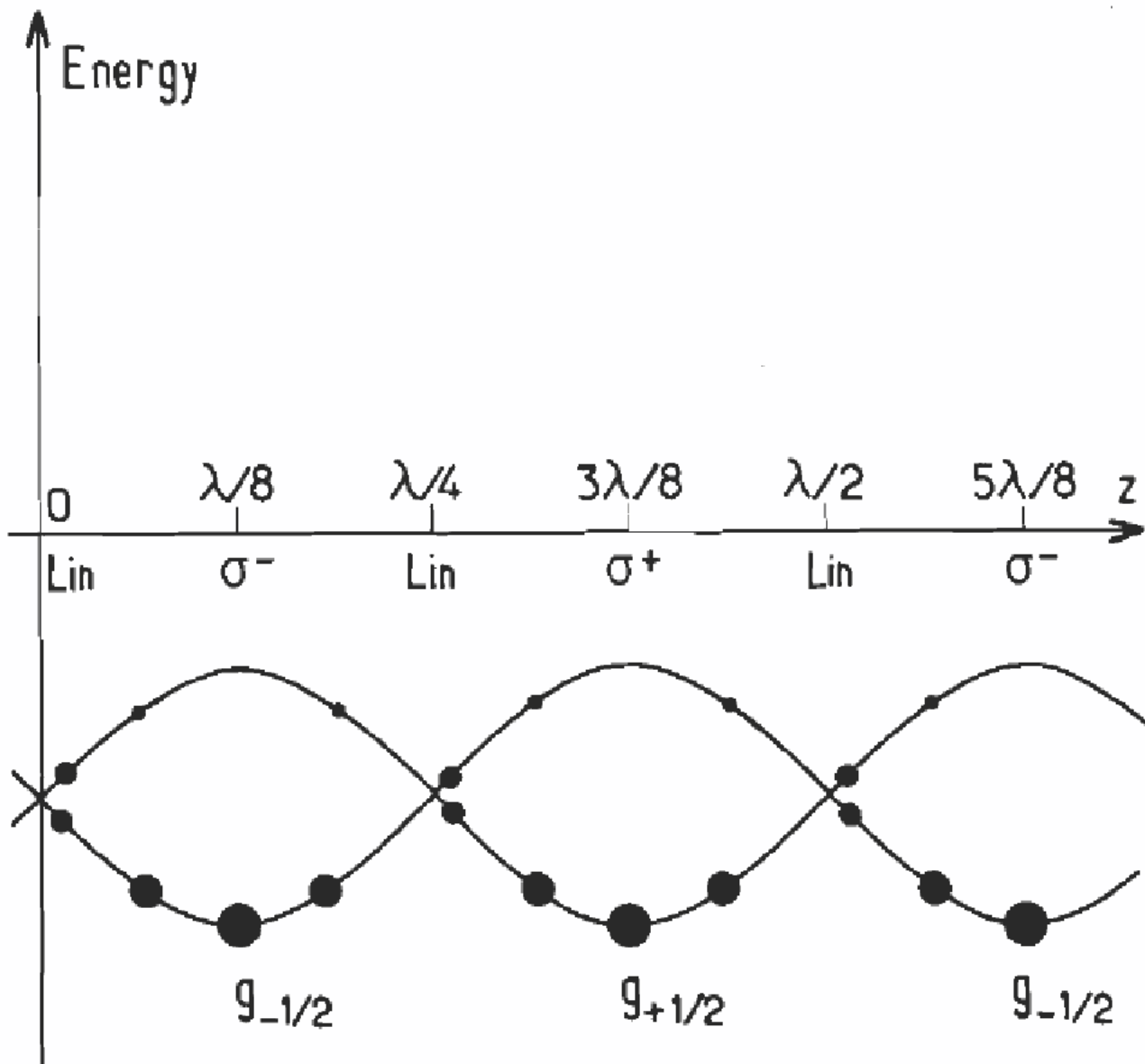


Fig. 3. Light-shifted energies and steady-state populations (represented by filled circles) for a  $J_g = 1/2$  ground state in the lin  $\perp$  lin configuration and for negative detuning. The lowest sublevel, having the largest negative light shift, is also the most populated one.

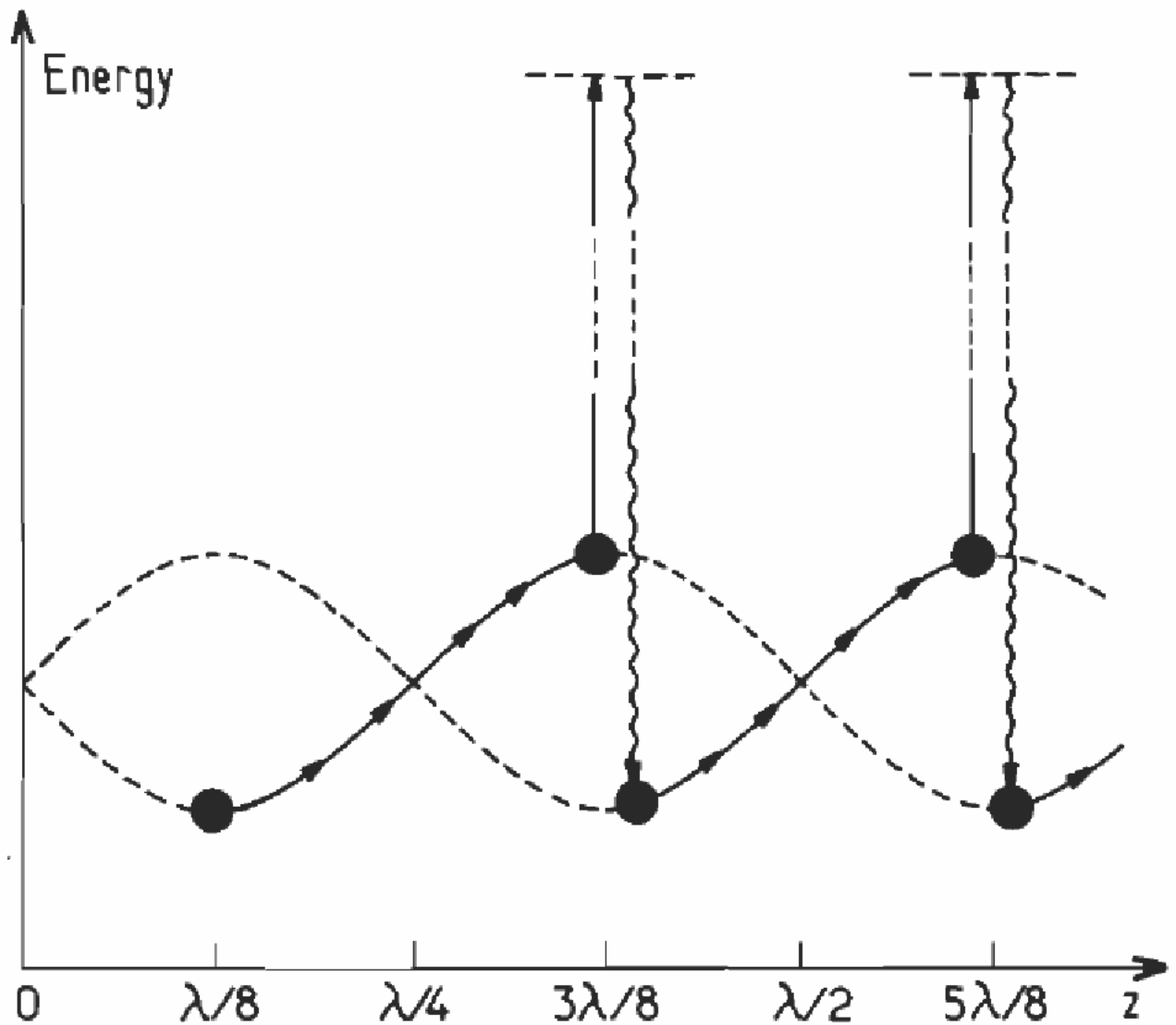


Fig. 4. Atomic Sisyphus effect in the lin  $\perp$  lin configuration. Because of the time lag  $\tau_p$  due to optical pumping, the atom sees on the average more uphill parts than downhill ones. The velocity of the atom represented here is such that  $v\tau_p \sim \lambda$ , in which case the atom travels over  $\lambda$  in a relaxation time  $\tau_p$ . The cooling force is then close to its maximal value.

# Sub-recoil cooling

Proof of impossibility

initial energy

$$\frac{1}{2} m \vec{v}^2$$

after abs.

$$\frac{1}{2} (m\vec{v} + \hbar \vec{k}_L)^2$$

after em.

$$\frac{1}{2} (m\vec{v} + \hbar \vec{k}_L - \hbar \vec{k}_R)^2$$

$$\Delta E = 2 E_{\text{rec}} + \hbar (\vec{k}_L - \vec{k}_R) \cdot \vec{v}$$

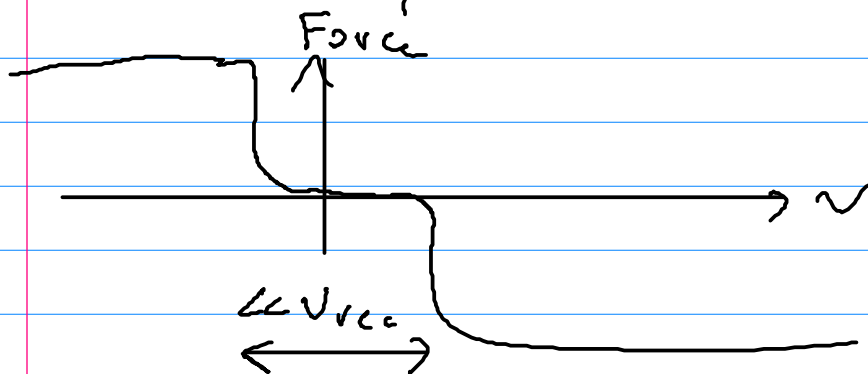
↑  
random

$$\overline{\Delta E} = 2 E_{\text{rec}} + \hbar \vec{k}_L \cdot \vec{v}$$

cooling requires  $\vec{k}_L$  antiparallel to  $\vec{v}$

$$\Rightarrow \overline{\Delta E} < 0 \text{ requires } v > v_{\text{rec}} = \frac{\hbar k_L}{m}$$

BUT: the last scattered photon may not be random



Velocity-Space  
Optical pumping

Practical implementations

- Raman cooling
- VSCPT